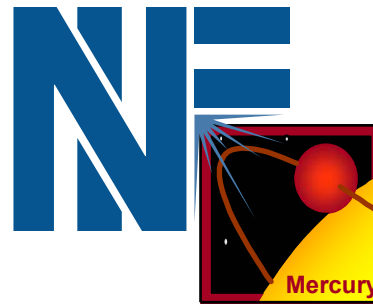


The Mercury Laser - Technical Progress Update



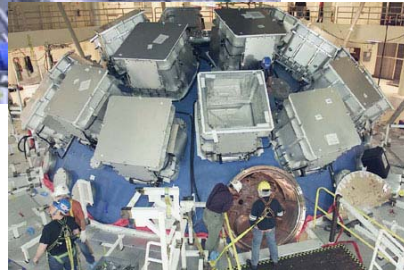
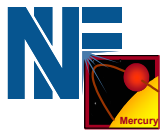
Camille Bibeau

**National Ignition Facility Directorate
Lawrence Livermore National Laboratory
Livermore, California 94550**

**Princeton Plasma Physics Laboratory
Princeton, New Jersey
October 27-28, 2004**

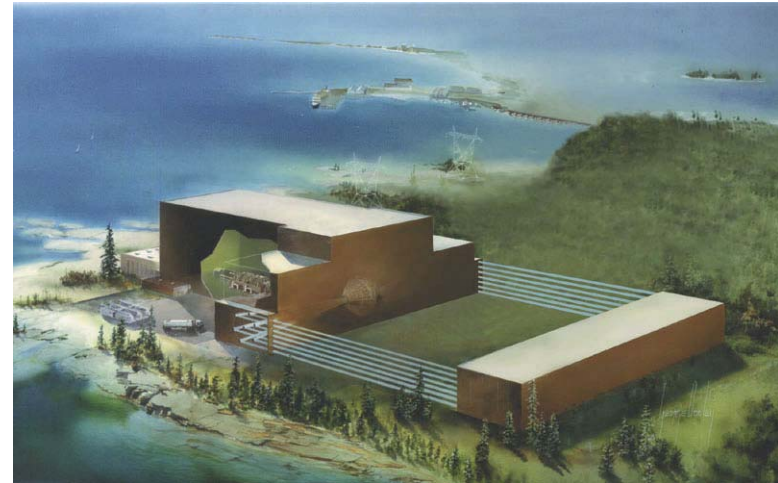
- **Project Overview**
 - Mercury Laser performance goals
 - International 100 J class systems
- **Laser architecture**
 - **Design**
- **System performance**
 - Pump diode arrays
 - Crystalline gain media
 - Gas cooled amplifiers
 - **Laser operations**
- **Upcoming activities**
 - **Frequency conversion**
 - Front end (Andy Bayramian: Poster)
 - Next generation design considerations (Ray Beach: Poster)

Mercury laser requirements are a melding of both NIF and IFE systems but at sub-scales



- Energy: 1.8 MJ
- Pulse shape: 3 ns shaped
- Bandwidth: 90 GHz; 1ω
- Wavelength: 0.35 μm

- Efficiency: 10%
- Reliability: $>10^9$ shots
- Cost: \$500/J for laser, \$0.05/W for diodes
- Repetition rate: 5-10 Hz

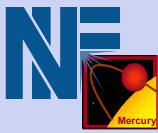


Mercury Laser incorporates:

Diodes, crystals, and gas cooling

- Energy: 100 J
- Pulse shape: 3 -10 ns
- Bandwidth: 150 GHz; 1ω
- Wavelength: 0.5-0.35 μm
- Efficiency: 10 % (w/o utilities)
- Reliability: $>10^8$ diode shots
- Cost: \$5/W for diodes
- Repetition rate: 10 Hz

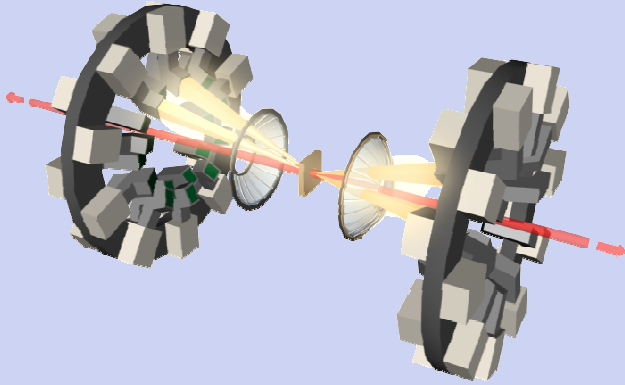
Many different architectural approaches are being considered for rep-rated 100J systems



Polaris - Germany

Dr. Joachim Hein

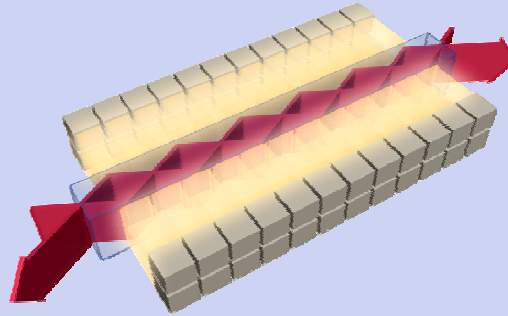
Water cooled, longitudinal pumped
Yb:Fluorophosphate disk



HALNA - Japan

Dr. Yasukazu Izawa

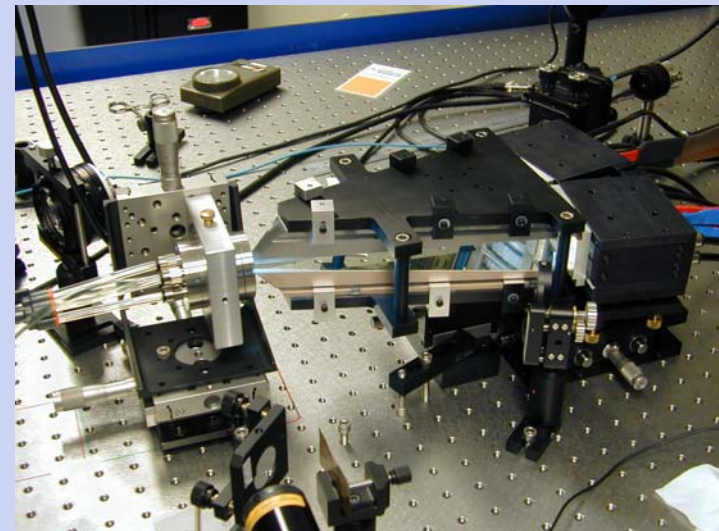
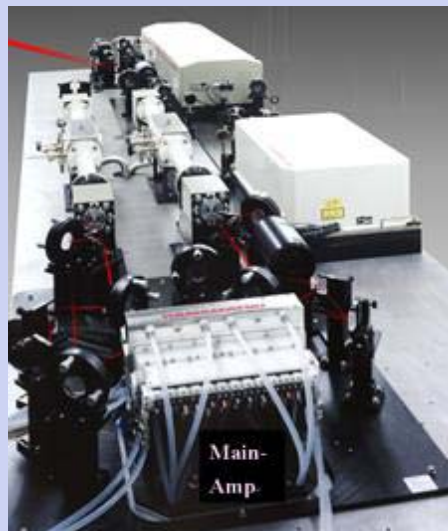
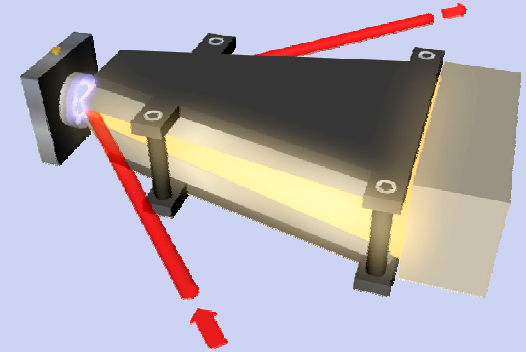
Water cooled, side pumped
Nd:Phosphate slab



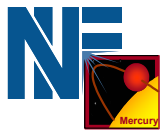
Lucia - France

Dr. Jean-Christophe Chanteloup

Water cooled, longitudinal pumped
Yb:YAG disk



Summary table of performance goals for high energy 1 μm DPSSL efforts around the world

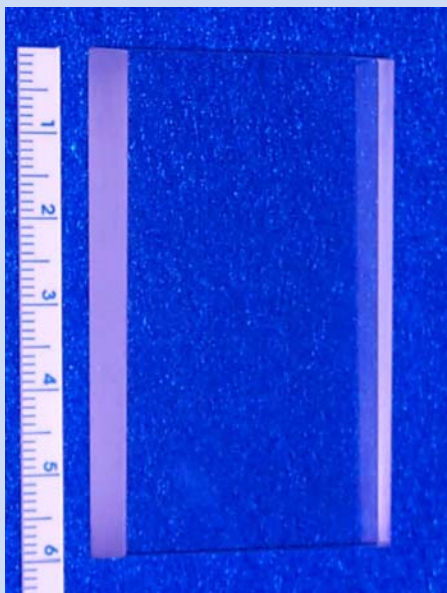


Project	Polaris Germany	HALNA Japan	Lucia France	Mercury United States
Application	High energy radiation source	IFE	Laser matter interaction	IFE
Gain Media	Yb:FP glass	Nd:phosphate glass	Yb:YAG and FP glass option	Yb:S-FAP
Energy	150 J	100 J	100 J	100 J
Rep-rate	0.1 Hz	10 Hz	10 Hz	10 Hz
Average Power	15 W	1 kW	1 kW	1 kW
Pulse length	150 fs	10 ns	1-10 ns	3-10 ns
Peak Power	1 PW	10 GW	10 GW	10 GW
Output beam	900 cm²	12 cm²	10 cm²	15 cm²
Optical Efficiency	10 %	20 %	20 %	20 %
Additional capabilities	Configured as a short pulse system	• >100 GHz	• 1 ps option	• Frequency conversion • 150 GHz • 10 ps option

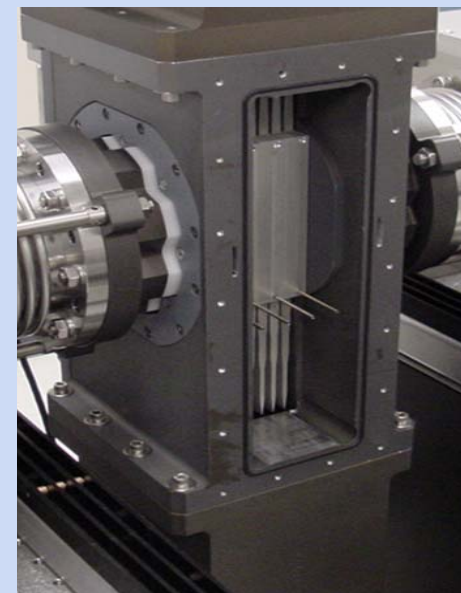
The Mercury Laser employs several key technologies



80 kW pump diode arrays

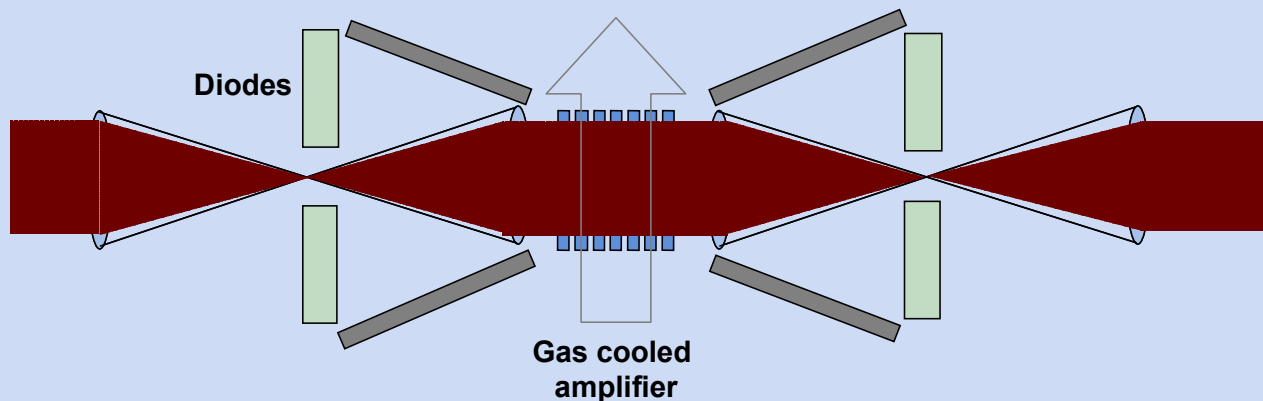


4x6 cm² Yb:S-FAP amplifier slabs

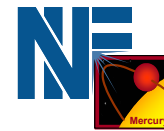


Mach 0.1 helium gas cooling

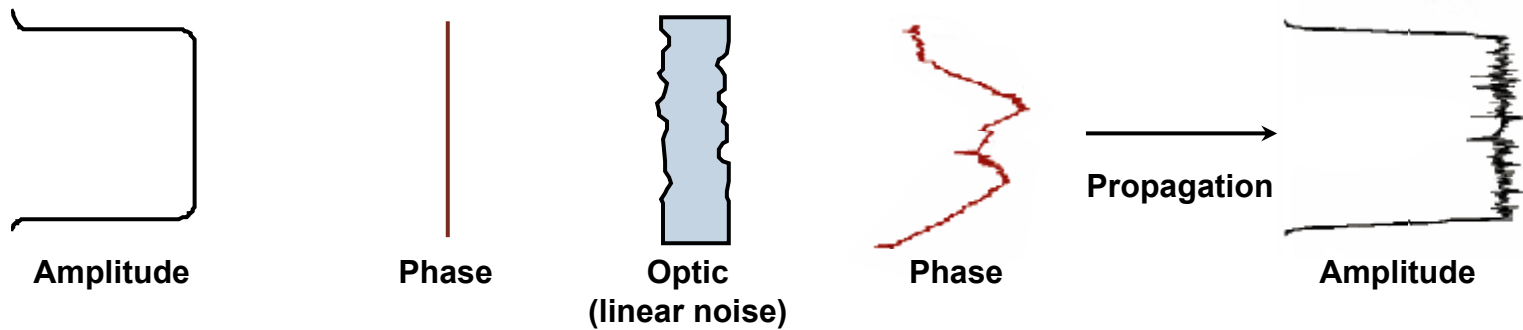
**Architecture:
closely-spaced
amplifier slabs**



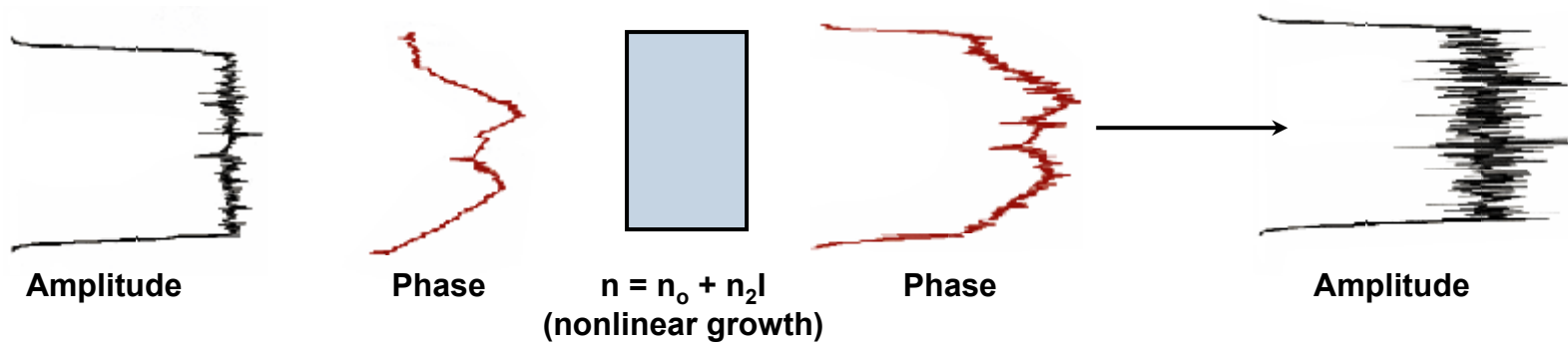
Architecture, optical specifications, and nonlinear propagation are addressed



Linear sources
of Amplitude variations



Nonlinear sources
of Amplitude variations

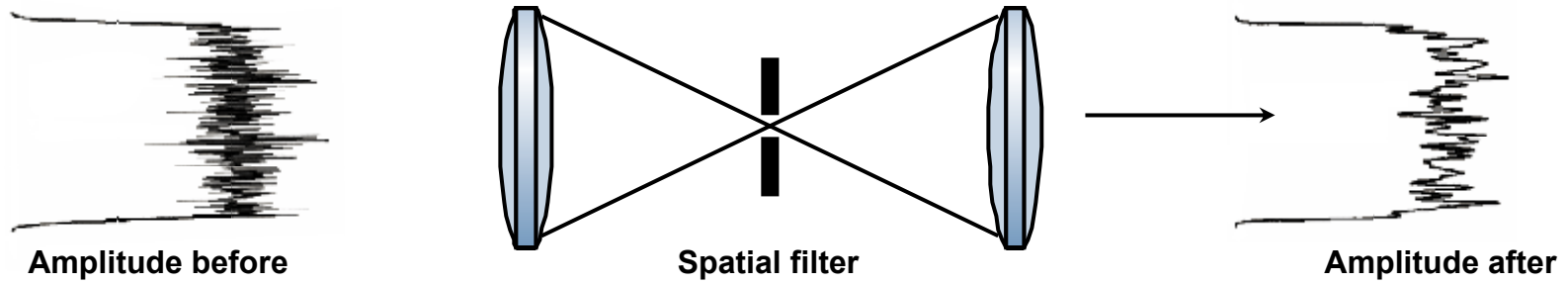


meters

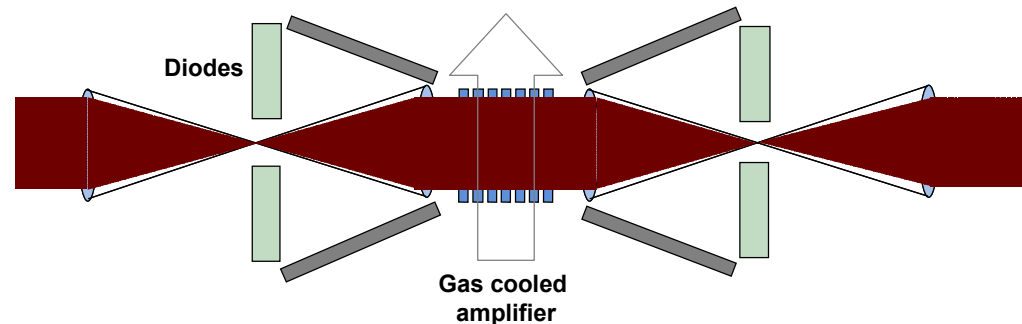
There are several ways to reduce nonlinear growth of beam modulation

Solutions:

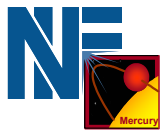
- Relay the location of near-field planes
- Filter fast growing spatial frequencies



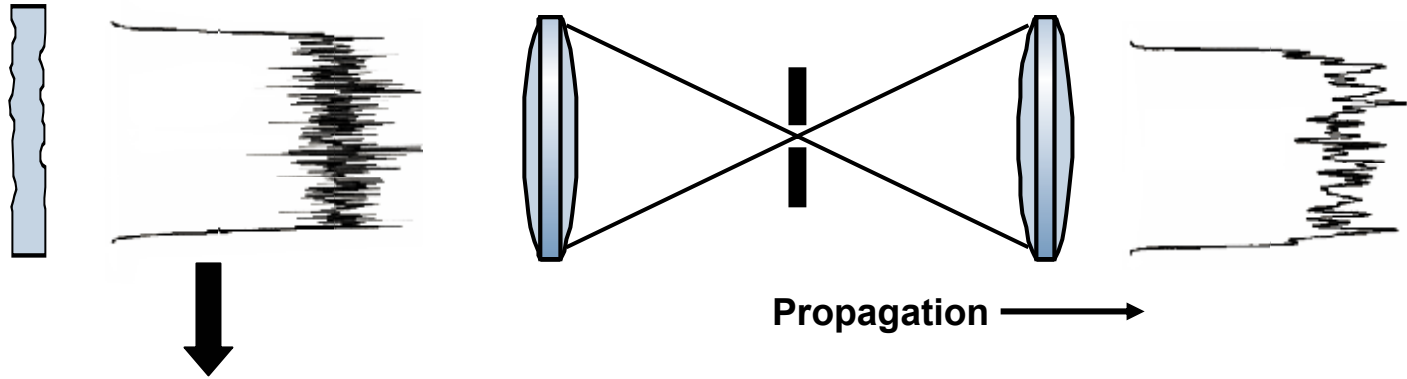
- Compact the optics near relay planes
- Minimize source terms through optical specifications



Spatial frequencies convert from phase to amplitude at different propagation distances



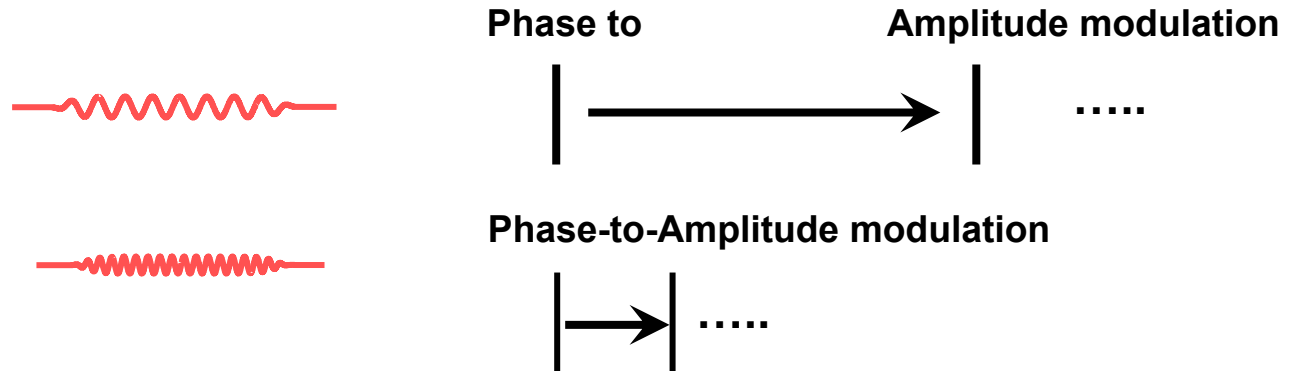
Amplitude Modulation



Phase aberration frequency decomposition



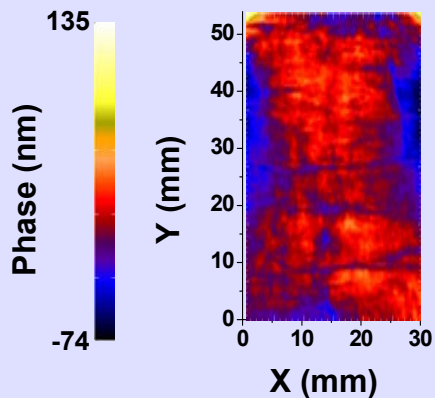
Cycling of phase to amplitude modulation on beam



When propagating, the highest frequency phase aberrations are the first to appear as amplitude modulation

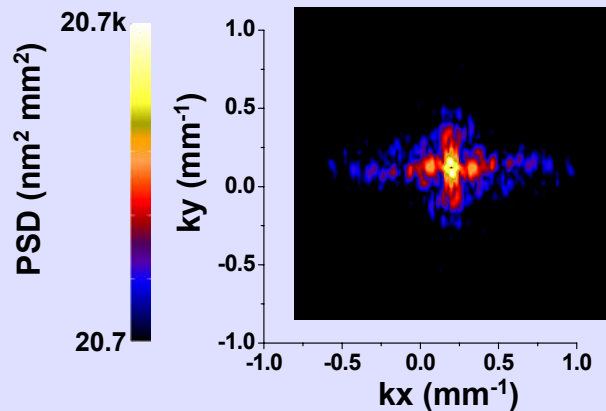
2-D Phase Map

$$\phi_{\text{Phase}}(x, y)$$



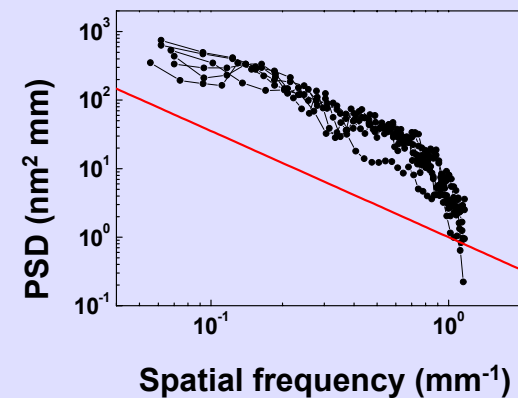
2D Power Spectral Density

$$\text{PSD}(v_x, v_y) = \frac{|\Phi(v_x, v_y)|^2}{\Delta v_x \Delta v_y}$$

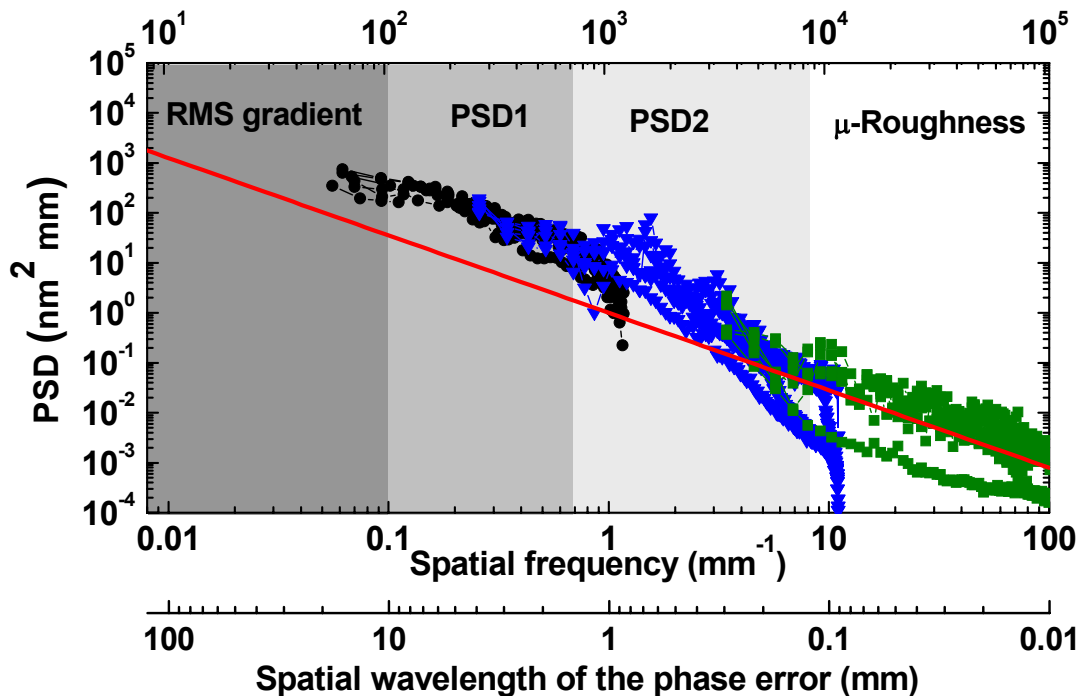


1 D Power Spectral Density

$$\text{PSD}(v_x) = \int \text{PSD}(v_x, v_y) dv_y$$



1w spectral half angle (μrad)



Optical specifications can drive:

**Beam
focusability**

**Pinholes
sizes**

**B-integral and
filamentation**

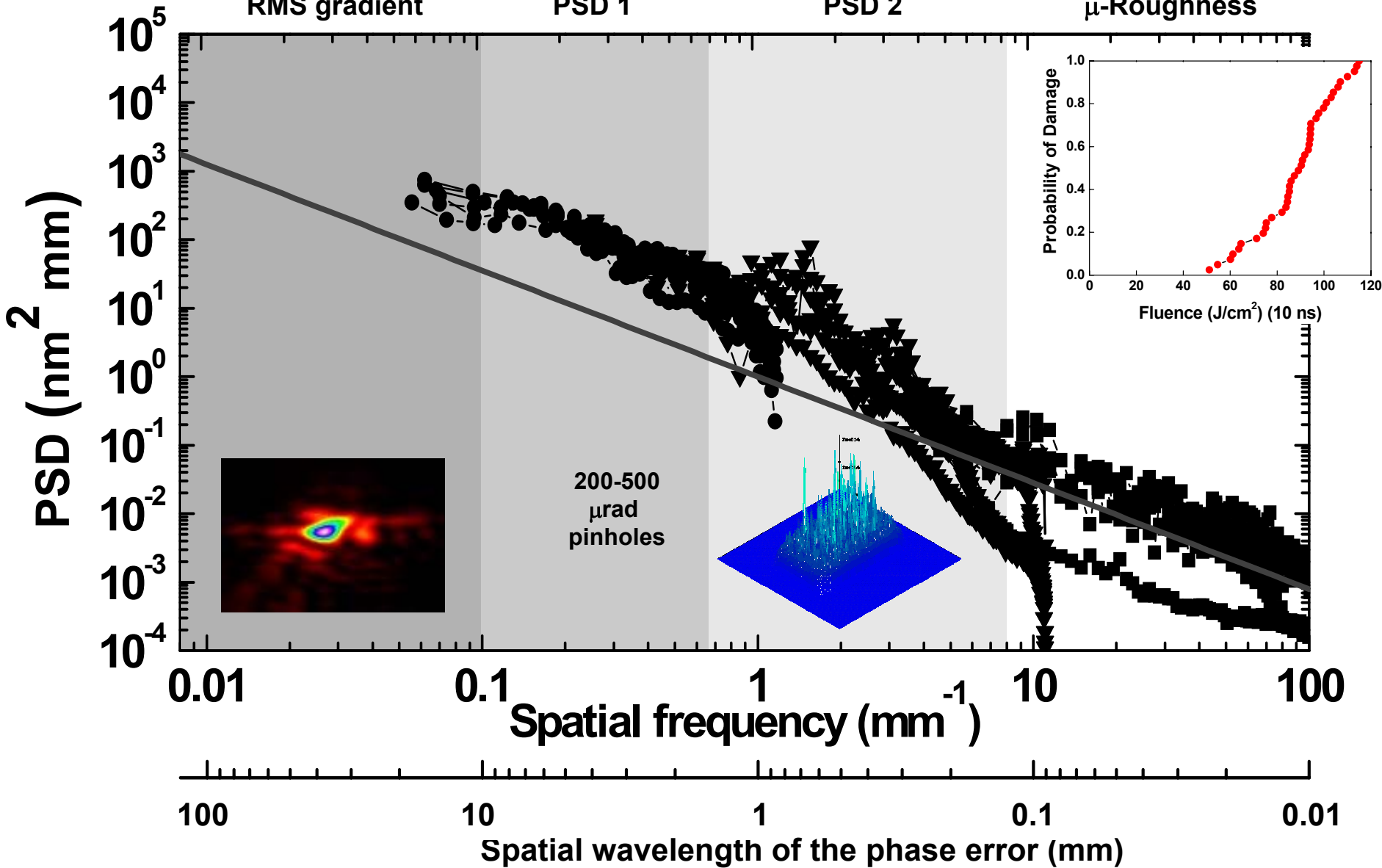
**Optical
lifetime**

RMS gradient

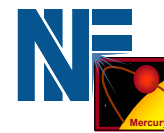
PSD 1

PSD 2

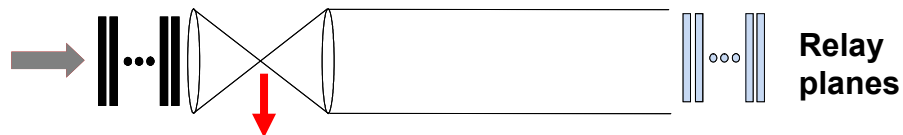
μ -Roughness



The closely spaced architecture reduces beam intensity on pinholes and optics

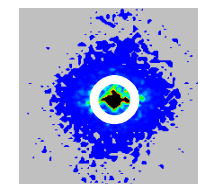
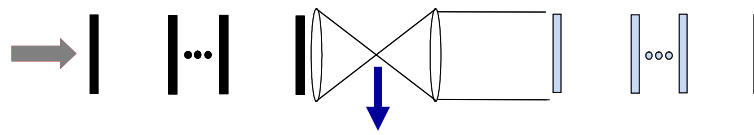


Closely spaced architecture

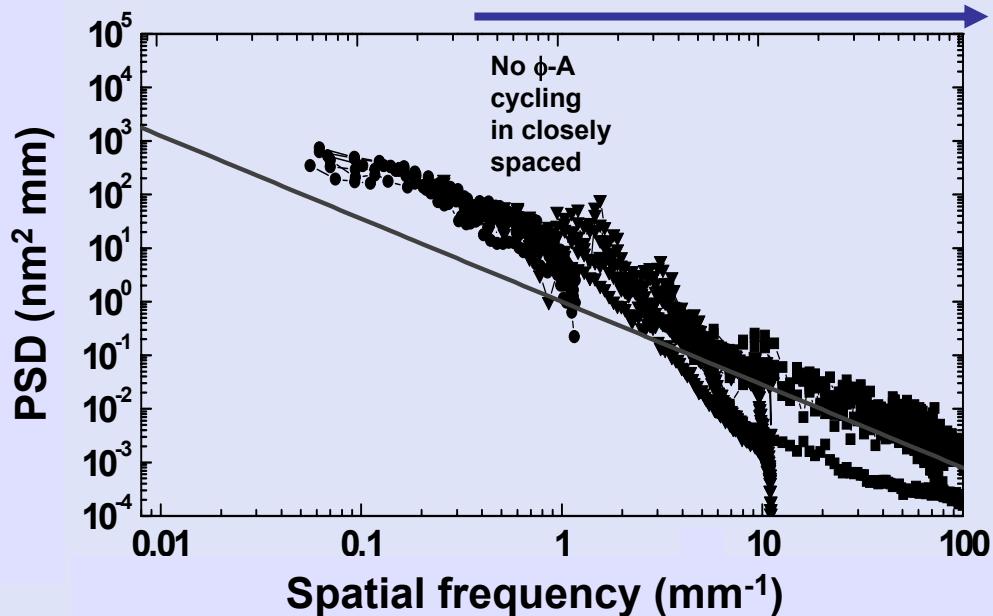


Pinhole intensity distributions

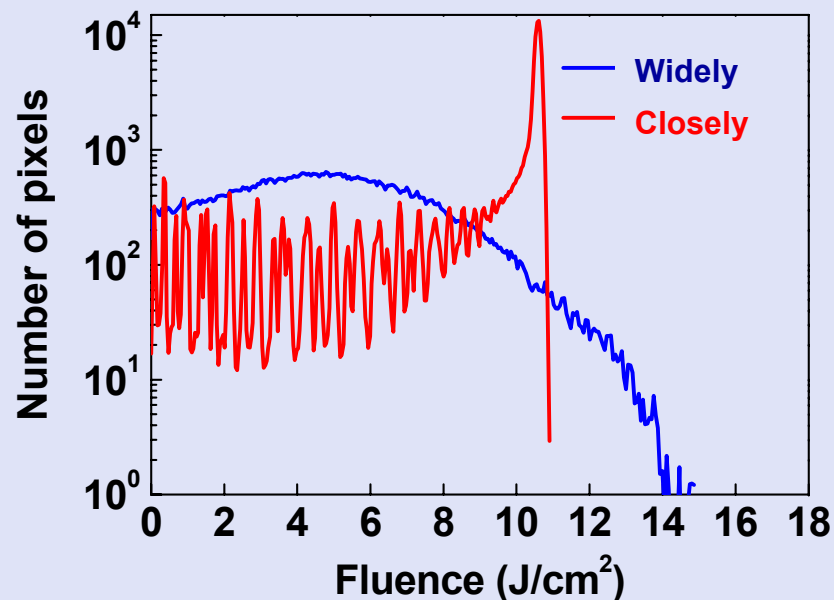
Widely spaced architecture



Noise term



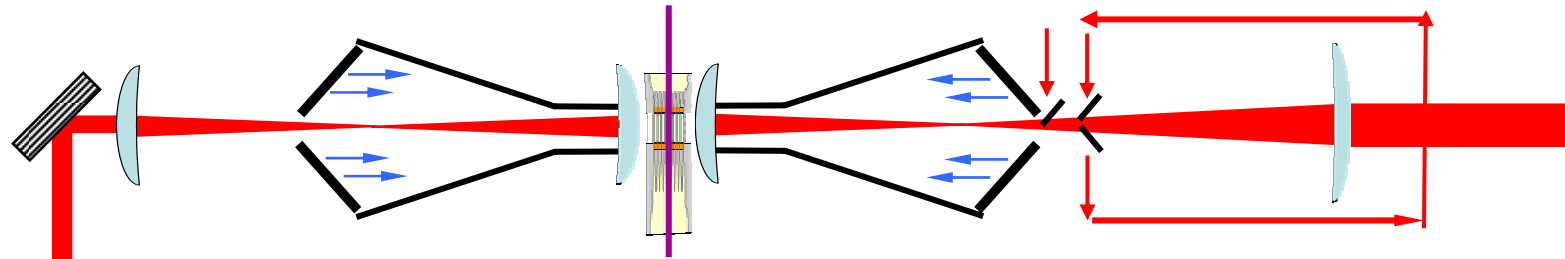
Near field histogram



On the basis of the reduced non-linear growth, Mercury was configured with closely spaced slabs

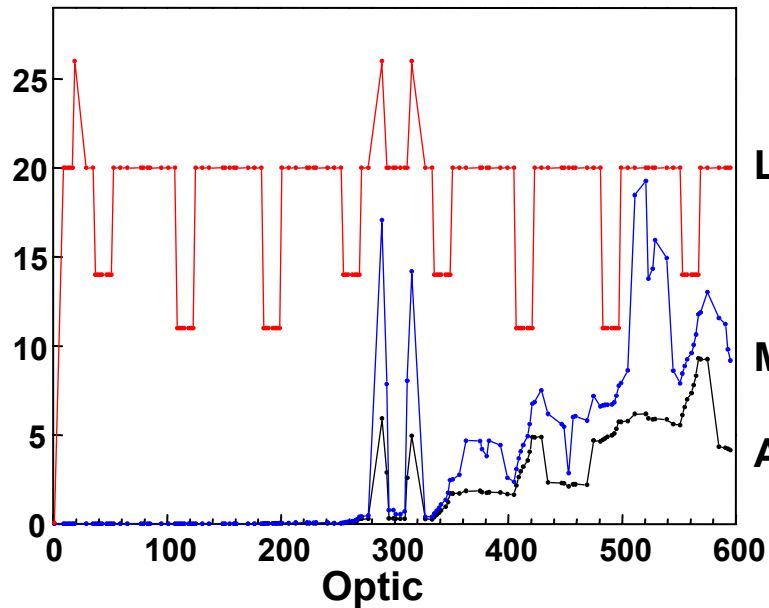
Diode Pump

S



Relay

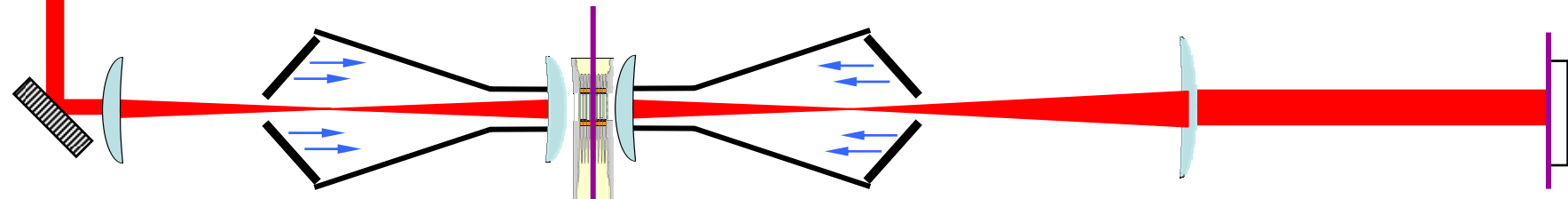
Fluence
 J/cm^2



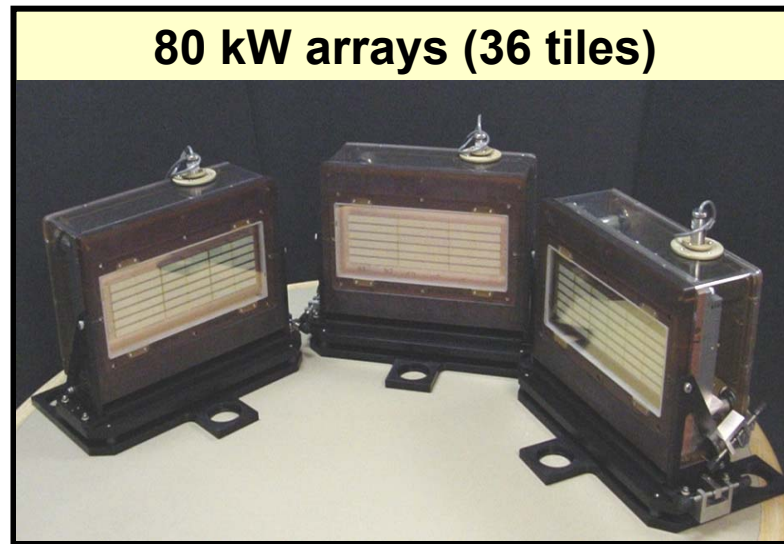
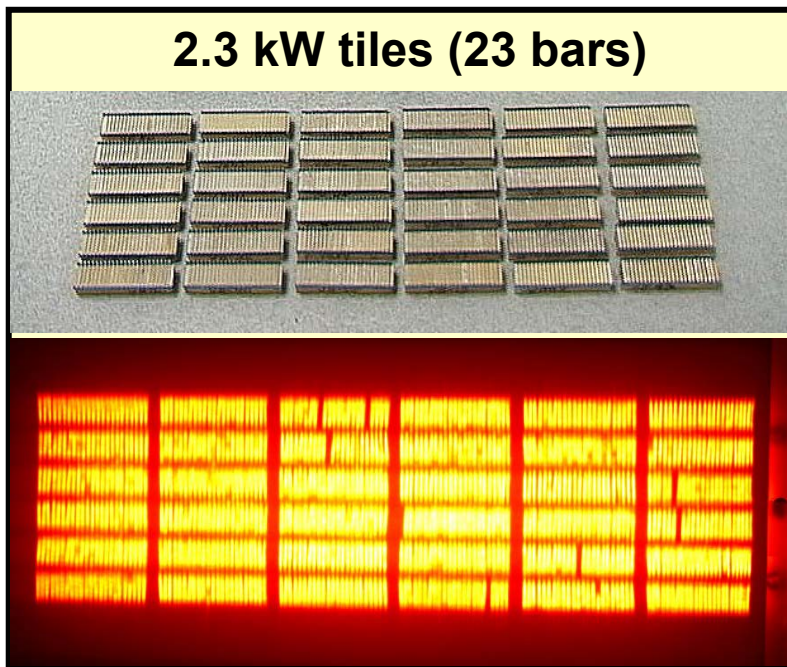
Limit

Maximum

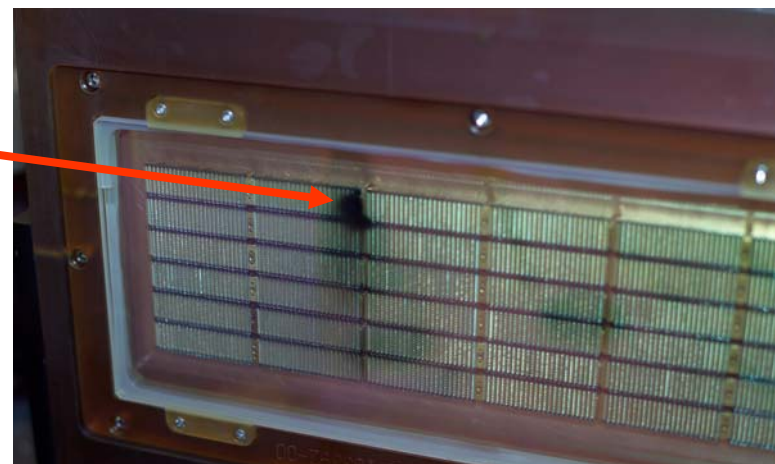
Average



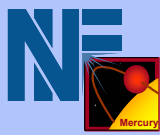
The Mercury Laser is currently running with all 8 diode arrays



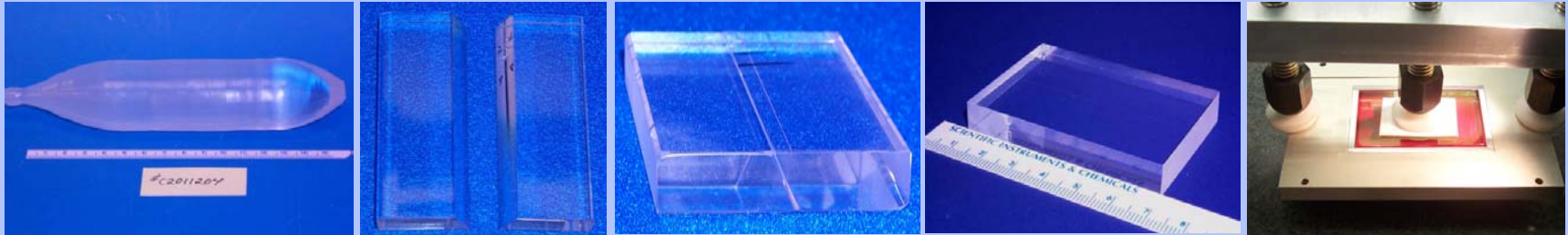
- During laser alignment a diode tile contact failed
- Analysis indicated contact fatigue occurred during installation
- New tooling is being developed
- Image monitoring analysis has been added



There are 23 slabs in fabrication to fully populate the amplifiers and provide 9 spares



Fabrication steps



G
4

S
6

B
1

S
9

A
2



Spare Slabs:



CZ Station 1 (small)



CZ Station 2 (small)

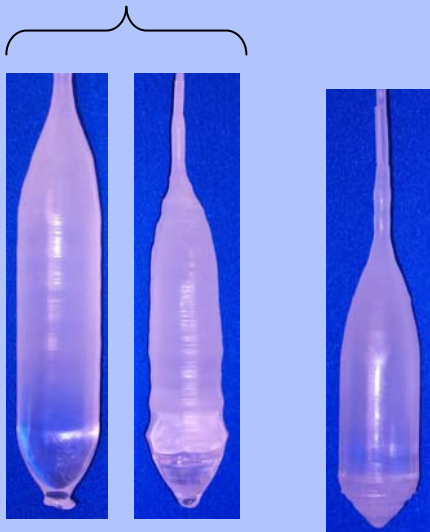


CZ Station 3 (large)



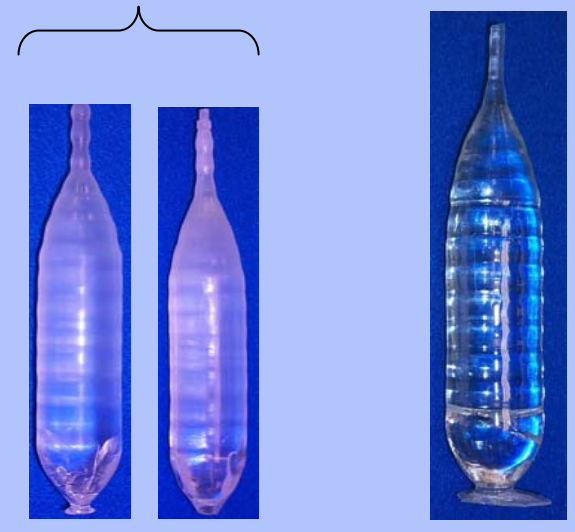
Before upgrade

After



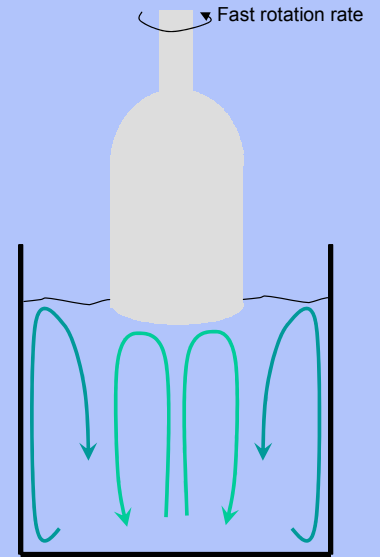
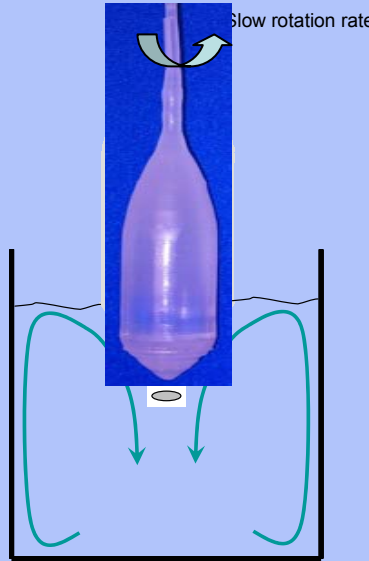
Before upgrade

After



**All three
LLNL crystal
growth stations
are operating**

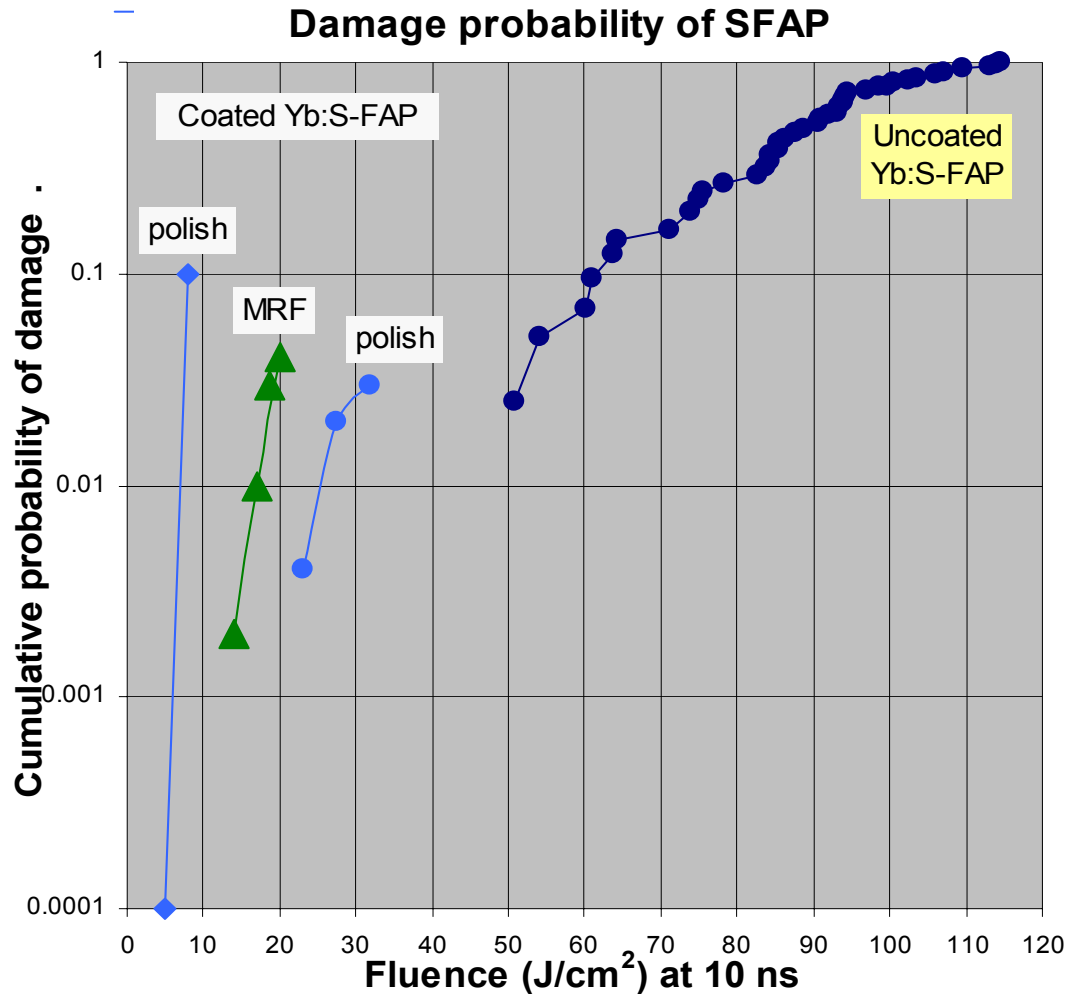
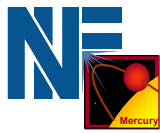
Eight 6.5 cm diameter boules have been were grown at



1st flat interface attempts



The impact of MRF finishing on 1ω laser damage threshold of coated S-FAP crystals is being evaluated

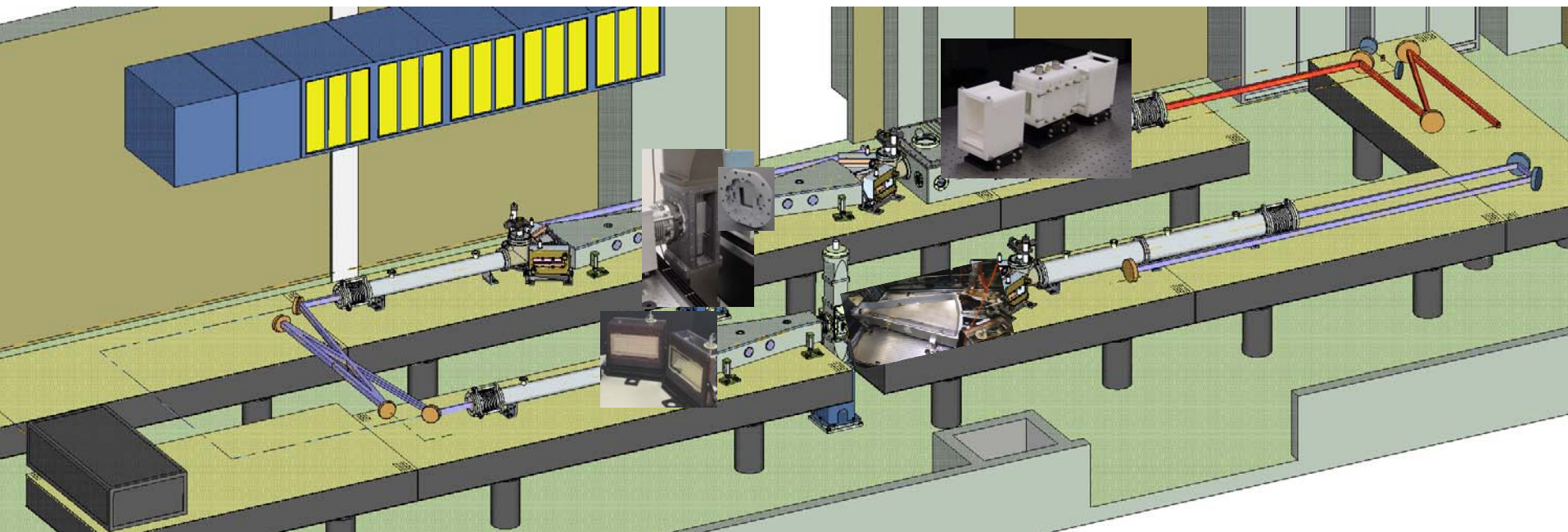
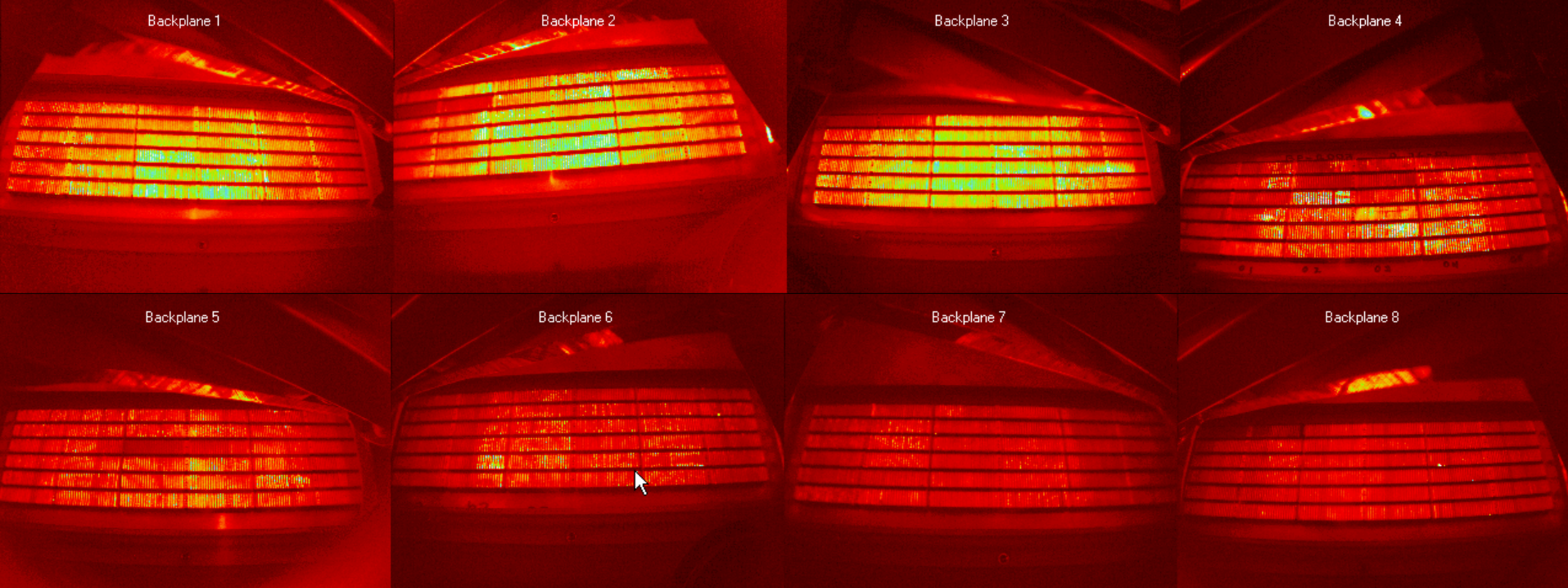


Damage Data Growth Experiments

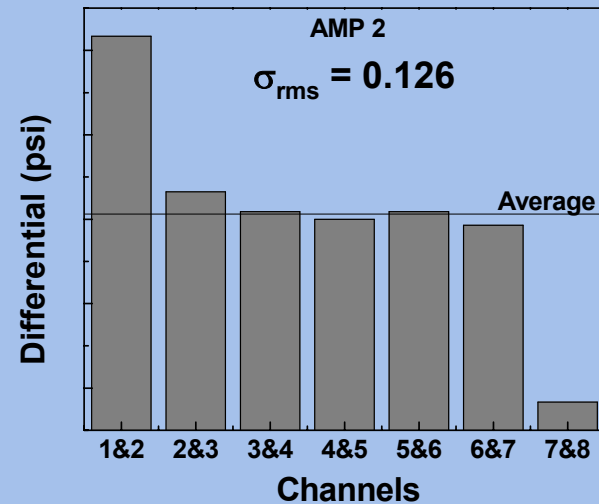
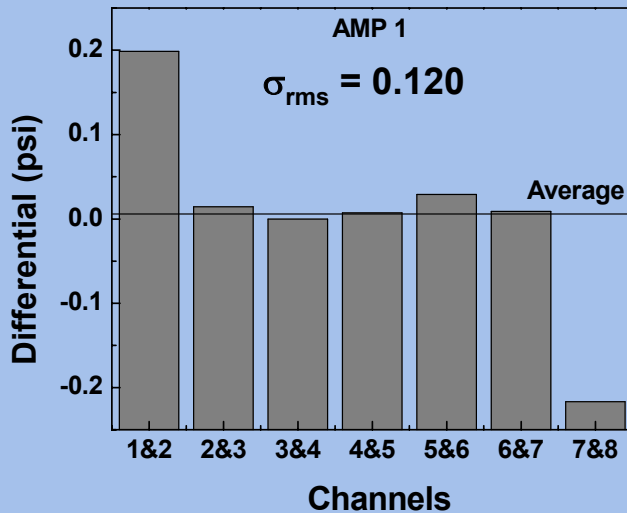
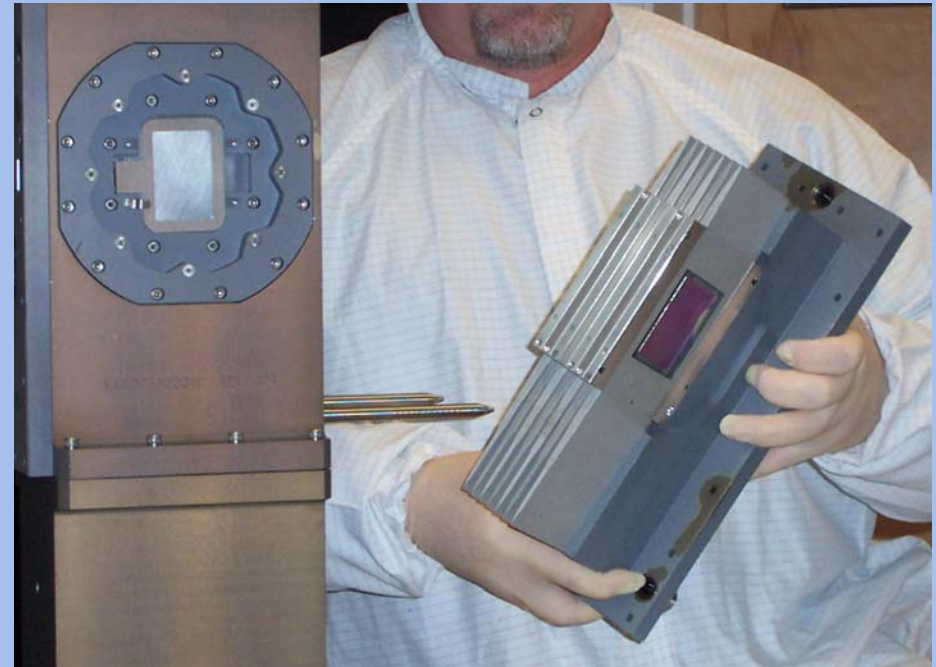
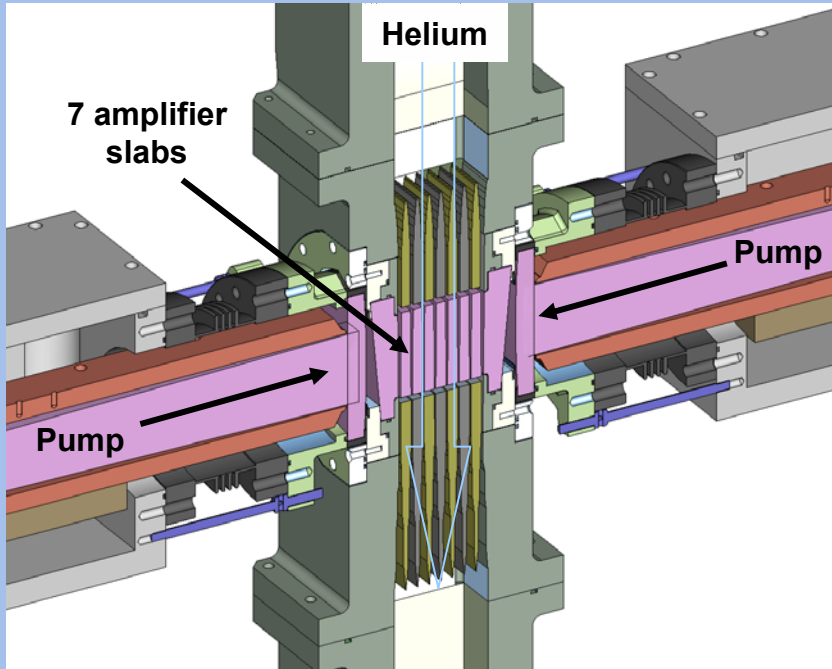
MRF
Initiation: 20 J/cm²
Additional sites: 20 J.cm²
Rapid degradation : 27.5J/cm²

Conventional Polish
Initiation: 20 J/cm²
Additional sites: 27 J.cm²
Rapid degradation: 30J/cm²

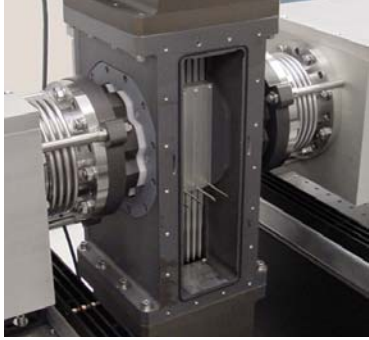




Both amplifiers have been deployed with helium gas cooling

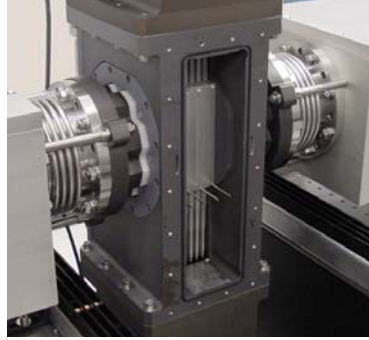


Amp 1

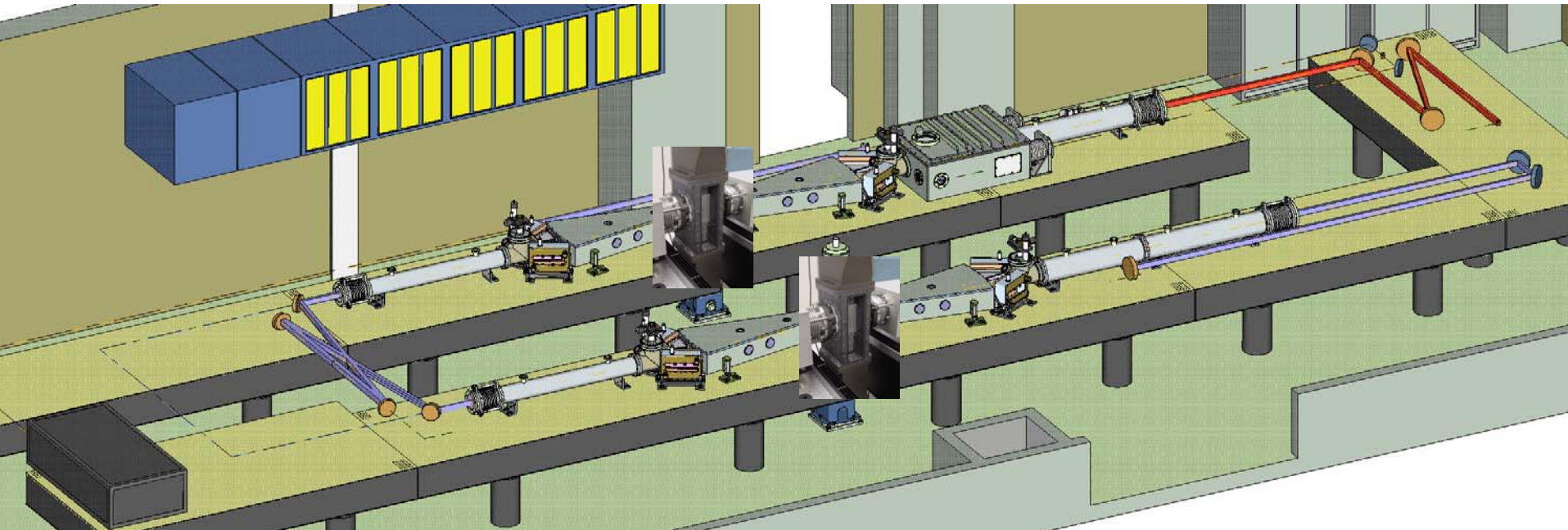
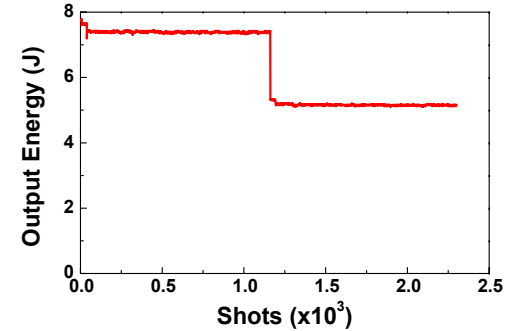
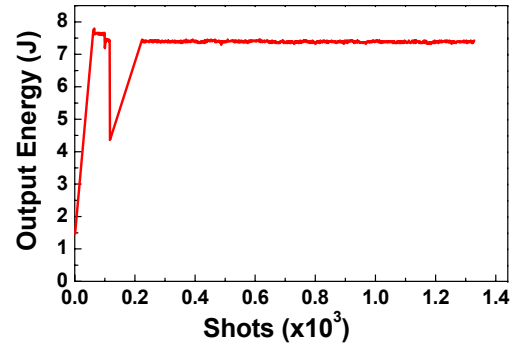


- 4 gain slabs
- 7 blanks
- 4 gain slabs
- 5 gain slabs

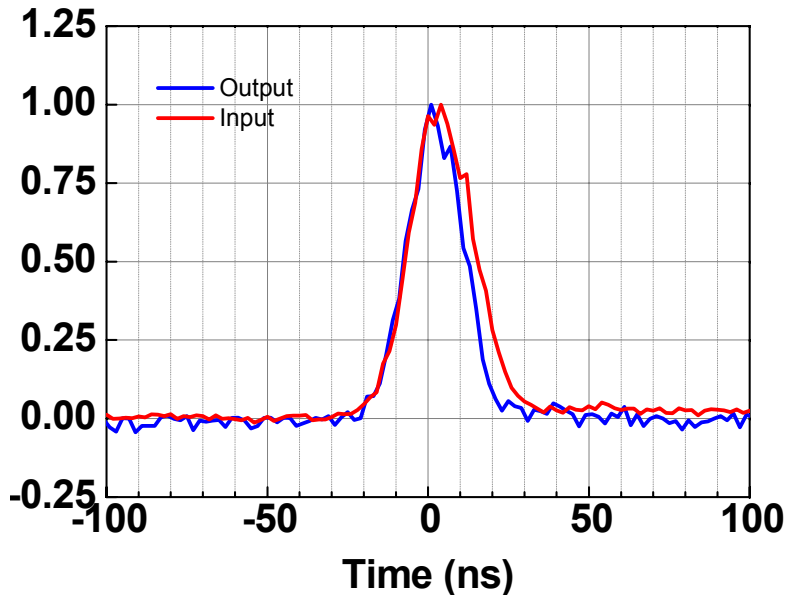
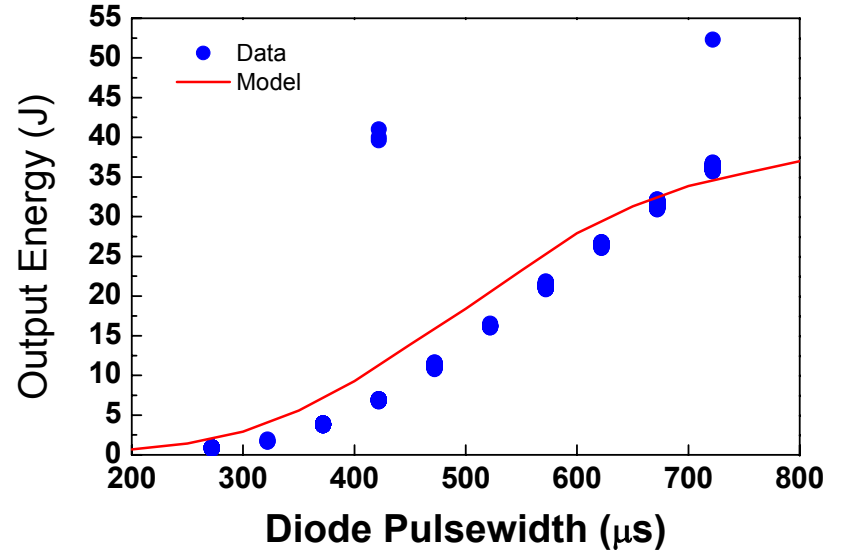
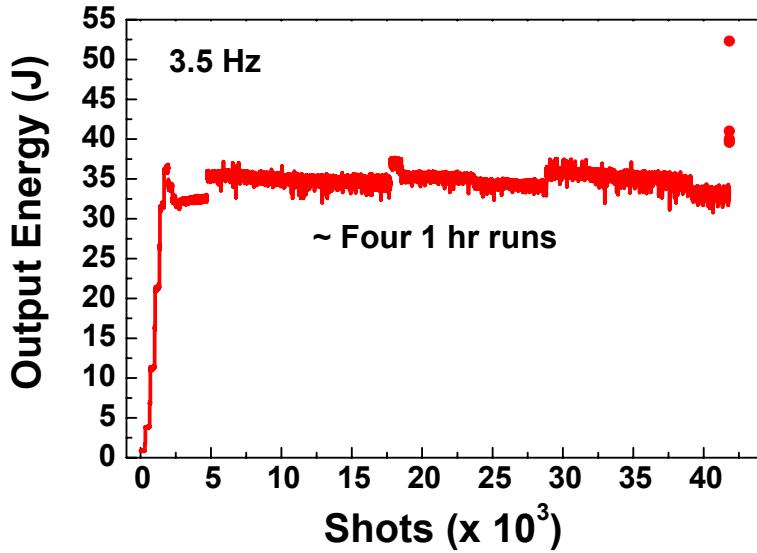
Amp 2



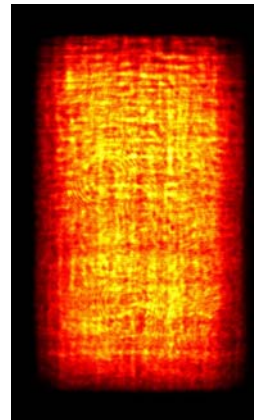
- 7 blanks
- 4 gain slabs
- 4 gain slabs
- 5 gain slabs (in progress)



Mercury 8 Slab Campaign Summary

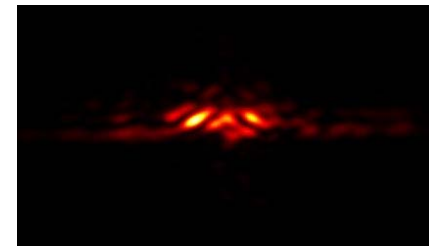


Output
Near field

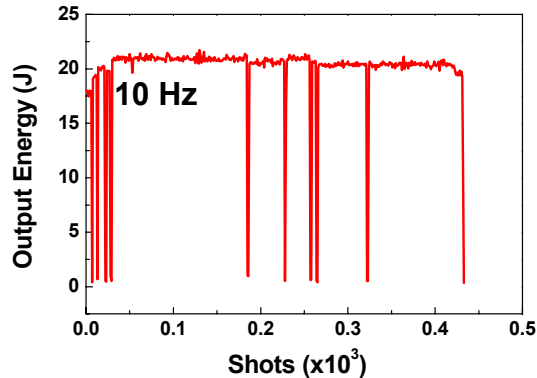
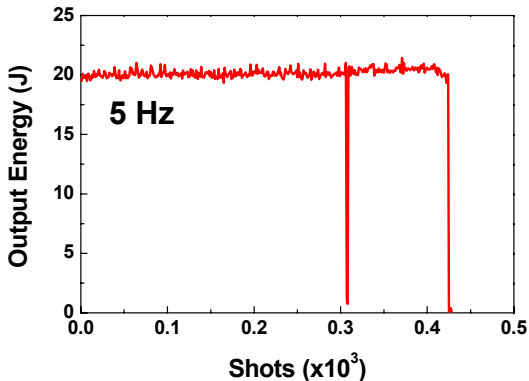


Output
Far Field

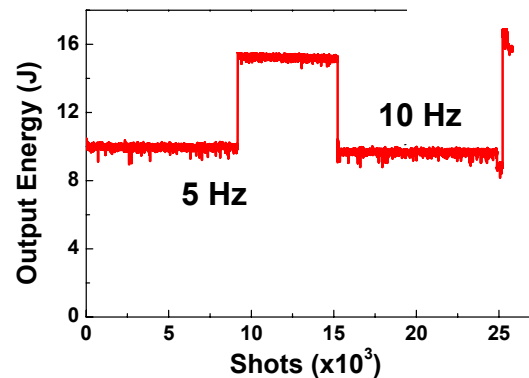
76% in 5X diffraction
limited beam



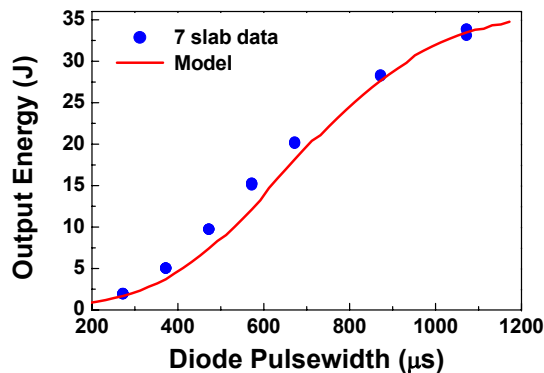
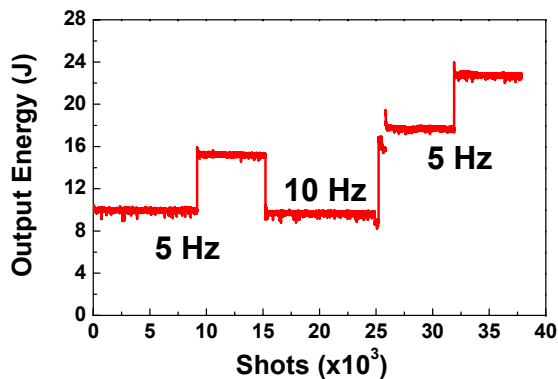
NRL: Dec 2002



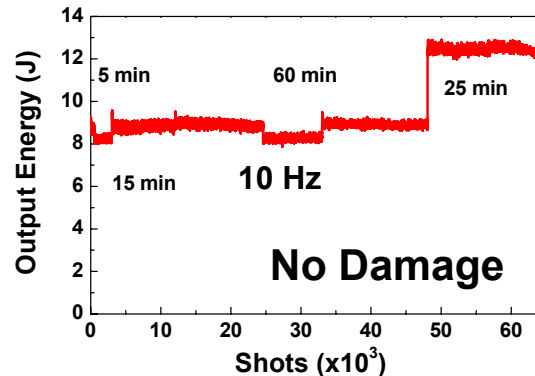
Sandia: Apr 2003



Wisconsin: Sep 2003

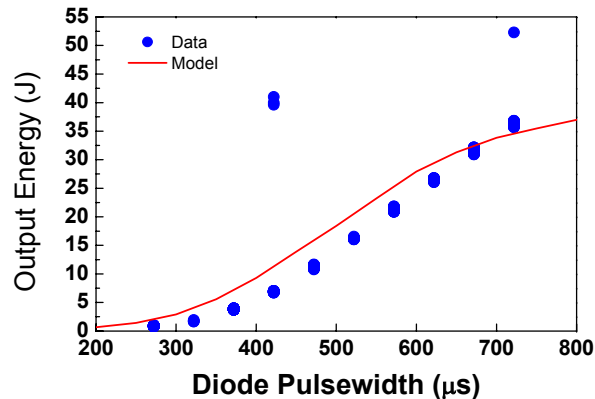
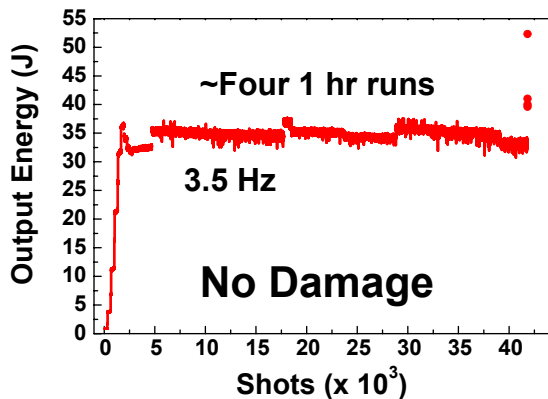


Atlanta: Feb 2004

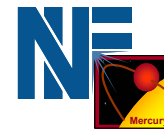


Princeton: Oct 2005

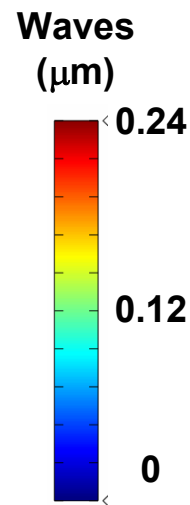
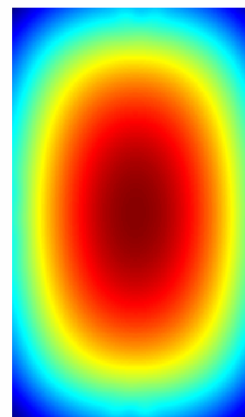
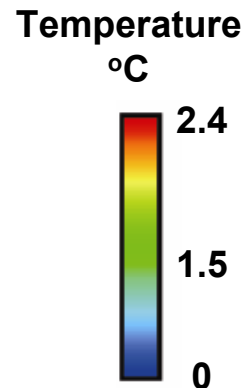
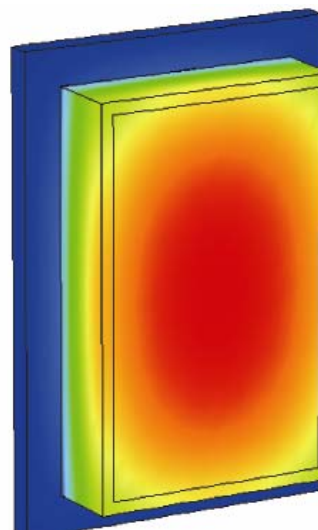
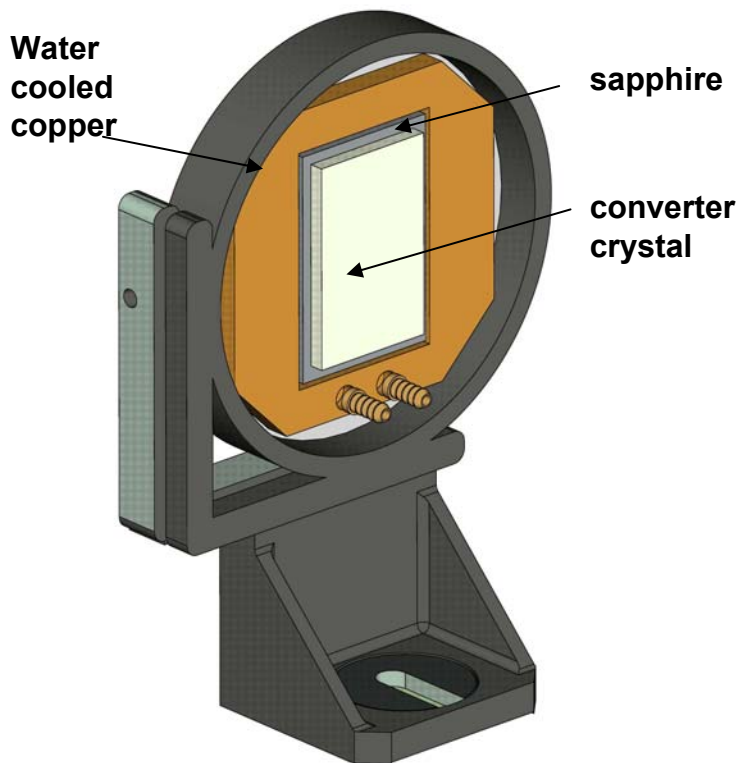
UCLA:
June 2005
2nd amp
build up



We have completed our thermal modeling of the frequency converter design

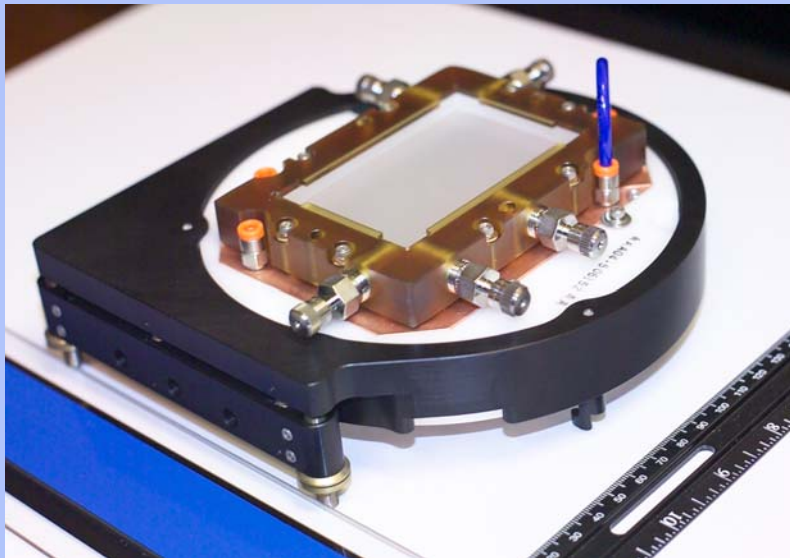


YCOB: 50J, 10Hz, 1.58 cm

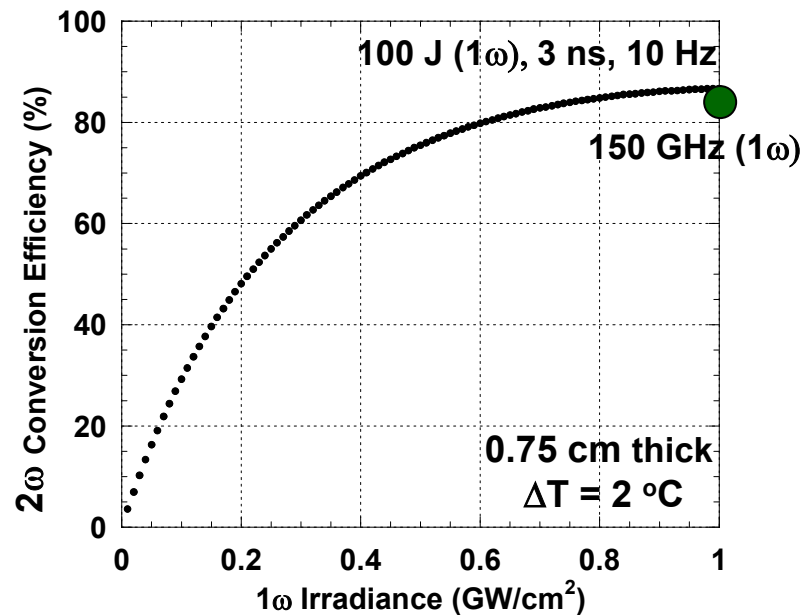
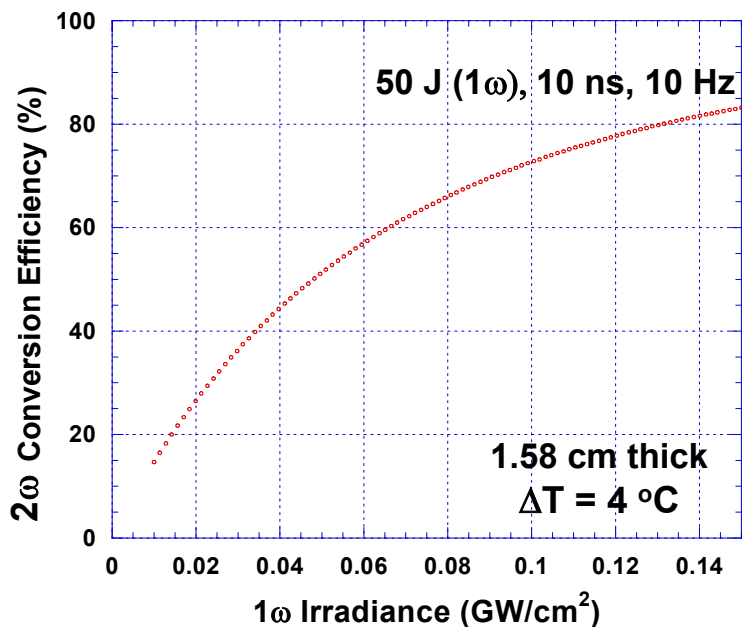
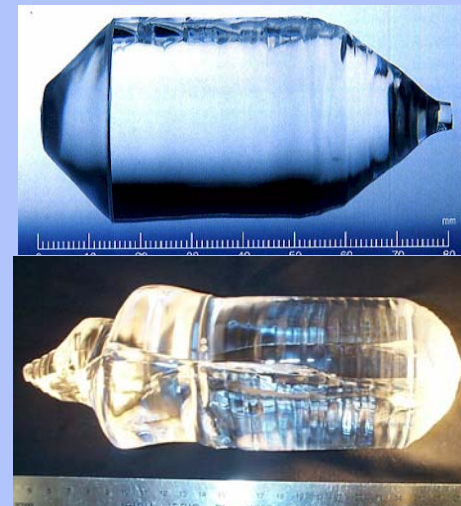


	Deff (pm/V)	Growth Achieved (dia. cm)	Angular Acceptance (mrad-cm)	Wavelength Acceptance (nm-cm)	Temperature Acceptance (°C-cm)
BBO	2.05	2	0.7	2.15	51
KDP	0.26	50+	1.25	19.7	11.3
DKDP	0.23	50+	1.34	5.2	~11
YCOB	1.1	8.5	1.22	1.15	40

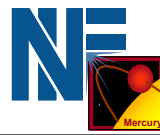
We have begun building the frequency conversion modules



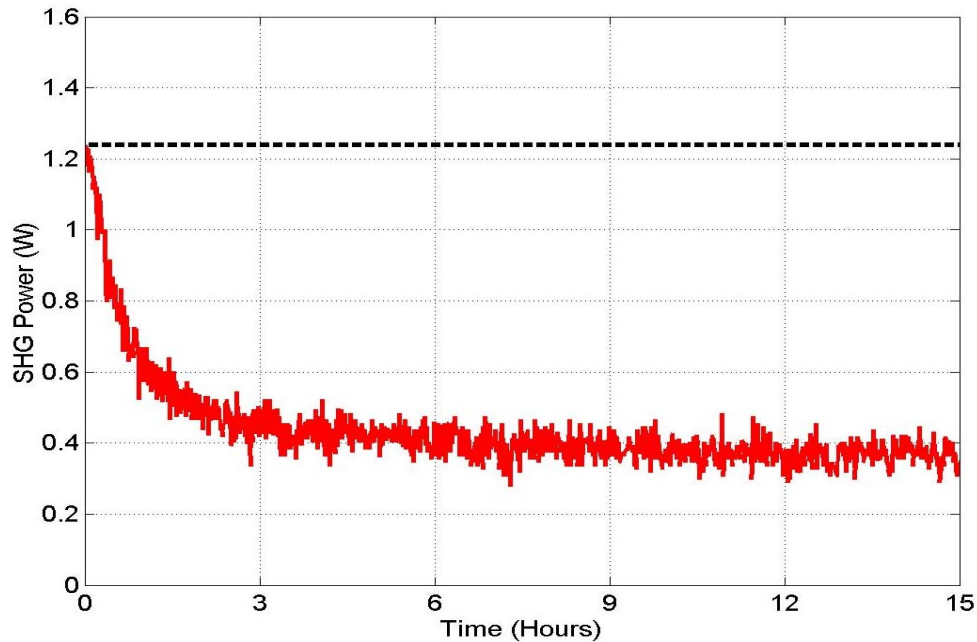
8 -10 cm long YCOB boules



YCOB lifetime experiments are being performed to confirm long term operation is possible without degradation in efficiency



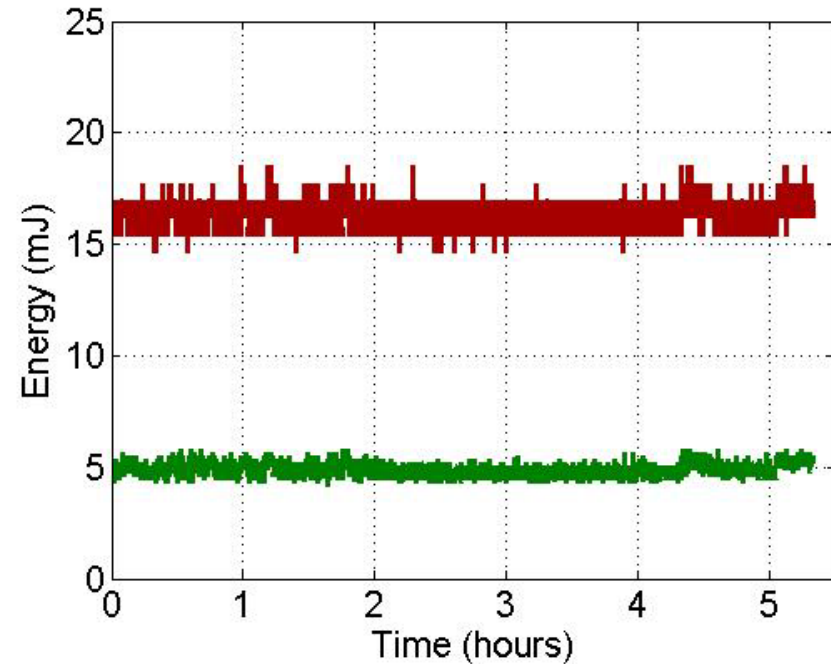
PPKTP



PPKTP exhibits photo-degradation:

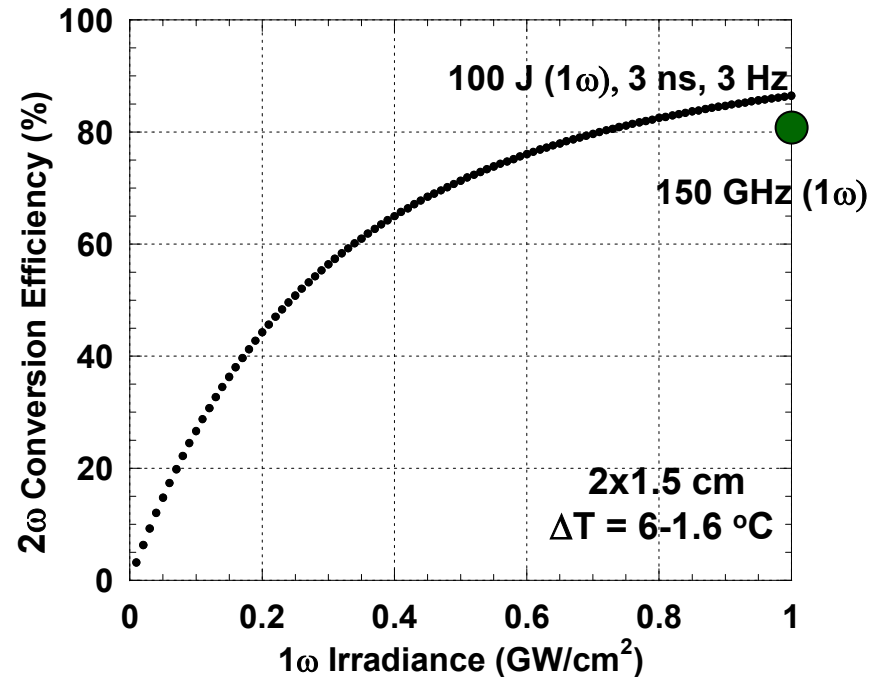
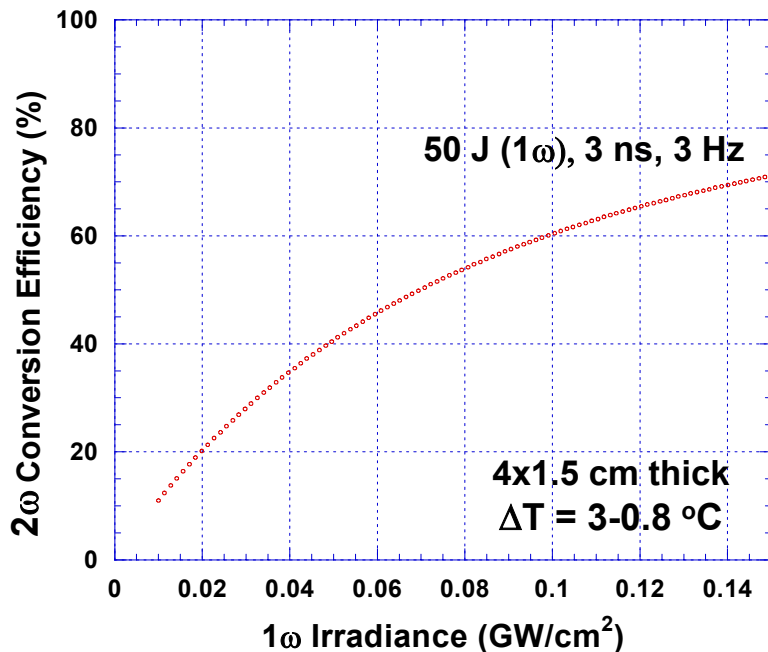
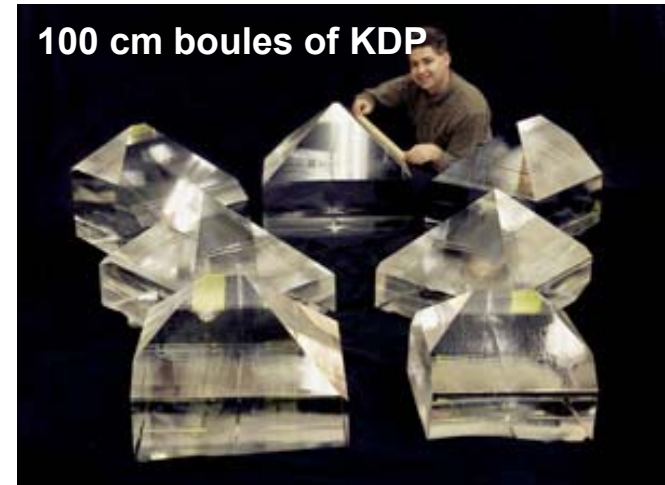
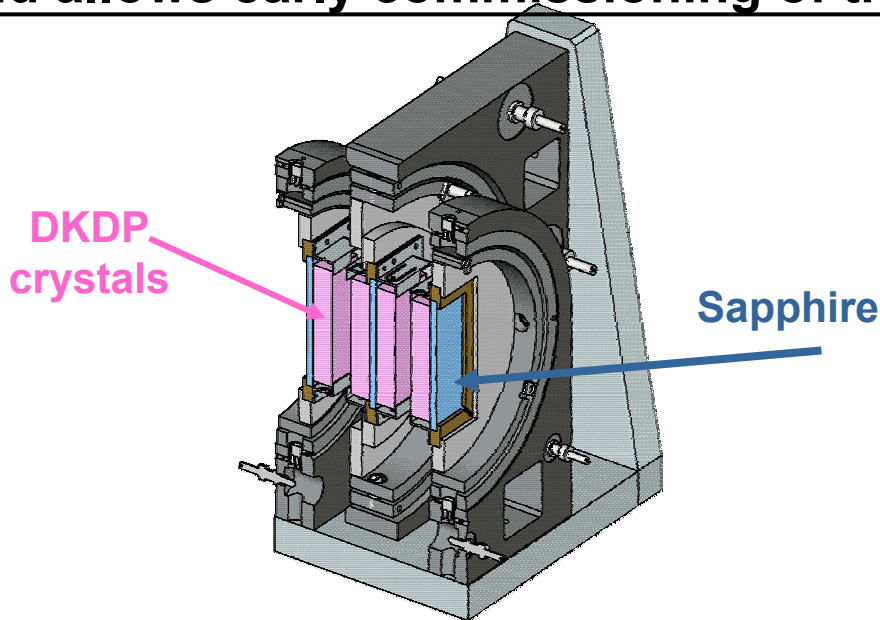
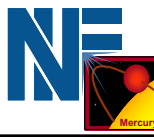
- color center formation (Ti^{4+} reduced to Ti^{3+})
- reduced conversion efficiency

YCOB



YCOB exhibits no degradation after 6 hours

DKDP is being pursued in parallel with YCOB and allows early commissioning of the cooling hardware



- **Project Overview (We are leveraging off of other DPSSL efforts)**
 - Mercury Laser performance goals
 - International 100 J class systems
- **Laser architecture (Architecture incorporates robustness)**
 - Design
 - Performance
- **System performance (System completed and several 1 hour runs accomplished)**
 - Diode arrays
 - Crystalline gain media
 - Gas cooled amplifiers
 - Laser operations
- **Upcoming activities and other topics**
 - Frequency conversion (Assembly of hardware begun)
 - Front End
 - IRE



IV



C

**Kathy Allen
Kathy Alviso
Paul Armstrong
Earl Ault
Monique Banuelos
Andy Bayramian
Ray Beach
Rob Campbell
Manny Carrillo
Chris Ebbers
Barry Freitas
Keith Kanz
Bob Kent
Tony Ladran
Dolores Lambert**

**Rod Lanning
Zhi Liao
Joe Menapace
Bill Molander
Noel Petersen
Greg Rogowski
Kathleen Schaffers
Ralph Speck
Chris Stolz
Steve Sutton
John Tassano
Steve Telford
Peter Thelin
Everett Utterback**

Laboratory for Laser Energetics

**CEA (Bordeaux)
Northrop-Grumman
Onyx Optics
Schott Glass Technologies
Spectra Physics
Quality Thin Films
Zygo
Photonic Crystals**

**Coherent
Directed Energy**